# ANALYSIS OF POLAR CAP ABSORPTION EVENTS II. TIME RELATION OF MAJOR FLARES AND RF EMISSIONS AT CENTIMETER WAVELTERIA RY

BY

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#### 1.0 RESULTS AND CONCLUSIONS

This study has been devoted to the analysis of 142 flares that were observed during the years 1954 through 1963, and assigned an importance of 3 or 3+ in the McMath-Hulbert Working List. We have studied the relationship of these flares to all known radio frequency emissions at 1500, 2000, 2800, 2980, 3000, and 3750 Mc/s, that occurred at the time of the flare. The area and magnetic classification of the sunspot groups on flare day, and the age of the associated plage region in solar rotations have been considered.

#### 1.1 TIME RELATION OF FLARES AND RF EMISSIONS

We find that for flares followed by an important PCA event the time of RF peak intensity coincided with or followed the time of the flare maximum intensity ( $\Delta t \ge 0$ ) in 76.5% of the cases; while  $\Delta t \le 0$  for 92% of the flares not followed by a PCA event.

#### 1.2 FLARES AND RF PEAK EMISSION

The RF peak intensity exceeded  $500 \times 10^{-22} \text{ W } (\text{m}^2\text{c/s})^{-1}$  for 87% of the flares that were followed by a PCA event while the flux was less than  $500 \times 10^{-22} \text{ W } (\text{m}^2\text{c/s})^{-1}$  for 76% of the flares not followed by a PCA event (including cases where the flare occurred during the normal observing time of at least one observatory and no flux was reported).

#### 1.3 FLARES AND SUNSPOT AREAS ON FLARE DAY

Of the major flares during 1954 through 1959, 71.8% occurred in sunspot groups that on flare day had an area ≥ 500 millionths

of the solar hemisphere while 92% of the flares that were followed by a PCA event had areas greater than 500 millionths.

#### 1.4 FLARES AND SUNSPOT MAGNETIC CLASS ON FLARE DAY

The magnetic classification of the associated sunspot group is not known on flare day for 21 spot groups. Ten of these flares occurred at either the east or west limb of the sun. Forty-seven per cent of the flares that occurred on days when the sunspot magnetic classification was known were classified as either  $\mathcal{E}$  or  $\beta \mathcal{E}$ . Only 3% of the  $\beta \mathcal{E}$  sunspots were followed by a PCA event. On the other hand, the flare-day magnetic classification is known for 32 of the PCA-flares and 6% occurred in either  $\mathcal{E}$  or  $\beta \mathcal{E}$  spots.

1.5 PERCENTAGE OF THE FLARES IN EACH CLASSIFICATION

The results of this analysis may be expressed as follows:

7% of all cases where △t ≥ 0 were followed by a PCA

90% of all cases where △t < 0 were not followed by a

PCA event

76.5% of the PCA flare △t≥0

92% of non PCA flares △t <0

87% of PCA flares the RF peak > 500 x  $10^{-22}$  W(m<sup>2</sup>c/s)<sup>-1</sup>

76% of non PCA flares, the RF peak  $< 500 \times 10^{-22} \text{W}(\text{m}^2\text{c/s})^{-1}$ 

72% of the flare-sunspots had an area ≥ 500 millionths of the solar hemisphere on flare day

92% of the PCA flare sunspots had an area > 500 millionths of the solar hemisphere on flare day

47% of the flare sunspots had a  $\mathcal V$  or  $\beta \mathcal V$  classification on flare day

69% of the PCA flare sunspots had a  $\chi$  or  $\beta \chi$  classification on flare day

NOTE: This study was based on a suggestion by Dr. Helen Dodson-Prince during a meeting of the Joint MSC-GSFC Working Group on Space Radiation. Dr. Prince suggested that the time difference between the maximum of the RF emissions and the optical H-alpha maximum should be investigated.

#### 2.0 CLASSIFICATION OF IMPORTANCE 3 AND 3+ FLARES

The 142 flares with importance 3 or 3+ in the McMath-Hulbert working list of flares (Table 1) were placed in three groups:

- (1) Flares followed by an important PCA event (Table 2).
- (2) Flares followed by a small or very small PCA event (Table 3).
- (3) Flares not followed by a known PCA event (Table 4).

#### 2.1 ALL FLARES OF IMPORTANCE 3 AND 3+

Table 1 lists all flares of importance 3 or 3+ that occurred during the years 1954 through 1963.

We have included the McMath Plage Number and plage age in rotations, the Mt. Wilson sunspot number, the Mt. Wilson magnetic classification, the sunspot area based on the Royal Greenwich Observatory photoheliographic results, and the Zurich classification are given for flare day when the values are available.

#### 2.2 IMPORTANCE 3 AND 3+ FLARES FOLLOWED BY A PCA EVENT

Table 2 lists all flares of importance 3 or 3+ that were followed by an important PCA event (from Table A-1). This table includes the flare, PCA, and short wave fade data. It also includes all known spectral emissions of Type II and IV associated with the flare, and single frequency data for frequencies 1500, 2000, 2800, 2980, 3000, and 3750 Mc/s. We have also included the value of the RF burst energy above 10% of the peak flux in units of 10<sup>-18</sup> joules (m<sup>2</sup>c/s)<sup>-1</sup> for 2800

and 3750 Mc/s when a value is given in the NASA Program Apollo working paper No. 1193 (January 28, 1966). The time difference in minutes △t is the time of the RF burst peak minus the time of flare maximum intensity.

2.3 IMPORTANCE 3 AND 3+ FLARES FOLLOWED BY VERY SMALL PCA EVENTS

Table 3 lists the reported small PCA events that were not included in Table 2 and those that were derived by Gregory based on the use of high sensitivity vertical incidence back scatter sounding of the lower ionosphere at a frequency of 2.3 Mc/s.

2.4 IMPORTANCE 3 AND 3+ FLARES NOT FOLLOWED BY A KNOWN PCA EVENT

Finally Table 4 lists all importance 3 and 3+ flares that were not followed by a known PCA event. This table gives all of the reported radio emissions in the frequency range from 1500 to 3750 Mc/s range as well as short wave radio fadeouts and reported spectral emssions of Type II and Type IV.

The number of importance 3 and 3+ flares for each year and the number followed by PCA events is summarized in Table 1.1. This shows that approximately 27% of the flares were followed by a PCA event if the PCA's reported by Gregory only are not included.

For the purpose of our analysis of the difference ( $\Delta$ t) between the time of RF emission peak flux and the time of the flares maximum intensity we have grouped the small PCA events with the flares not followed by PCA's.

	No. of Flares Imp. 3, 3+	With PCA	Small PCA	PCA Gregory Only
1955	4	0	0	
1956	21	2	0	
1957	32	n	1	
1958	20	6	0	
1959	32	6	1	
1960	19	7	1	7
1961	10	5	0	
1962	2	0	0	
1963	2	1	0	
TOTAL	142	38	3	7

TABLE 1.1 Number of Importance 3 and 3+ Flares and Associated PCA Events

# 3.0 RELATIONSHIP OF FLARE MAXIMUM INTENSITY AND PEAK RADIO FLUX INTENSITY

#### 3.1 FLARES FOLLOWED BY IMPORTANT PCA EVENTS

This list contains 38 PCA events, that with one exception, are included in one or more of the catalogues by Bailey, Modisette, Malitson, and Warwick and Haurwitz (Table A-2). Four of the events (Numbers 12, 17, 18, and 55, Table A-2, Flares Numbers 43, 47, 49, and 156, Table 1) were reported with a weak absorption and their inclusion in Table 2 rather than Table 3 is quite arbitrary.

All of the flares in this group were followed by a short wave radio fadeout and radio emissions at one or more of the frequencies between 1500 and 3750 Mc/s. The time of the maximum intensity of both the flare and the RF emission are reported for 34 of the events.

The time of the RF maximum occurred simultaneiously with or after the flare maximum ( $\Delta t \ge 0$ , where  $\Delta t =$ time of RF maximum minus the time of the flare maximum) for 26 of the events (76.5%) (Table 2.1). The eight cases where  $\Delta t < 0$  will be discussed in detail later when other factors will be considered.

abs. db.	Unknown	Positive	Zero	Negative	Total
≤ 3 B to	0 3	6	2	2	10 12
<b>≤</b> 6 >6 to <b>≤</b> 10	0	2	0	1	3
>10	1	9	1	2	13
TOTAL	4	23	3	8	38

TABLE 2.1 Relation Between the Intensity of the PCA Events and the Algebraic Sign of  $\triangle$  t

This table does not indicate any good correlation between the intensity of the PCA and algebraic sign of  $\Delta$ t.

On the other hand Table 2.2 shows that the peak RF flux exceeds 500 x  $10^{-22}$  W(m<sup>2</sup>c/s)<sup>-1</sup> for 86% of the PCA flares, and the  $\triangle t \ge 0$  for 80%.

A t Flux	Unknown	Positive	Zero	Neg <b>a</b> tiv <b>e</b>	Total
\$ 250	1	0	0	1	2
>250 to <500	ı	1	0	1	3
>500 to ≤1000	0	4	1	2	7
>1000 to ≥5000	1	13	1	3	18
>5000	1	5	1	1	8
TOTAL	4	23	3	8	38

TABLE 2.2 Relation Between the Peak Intensity of the RF Flux and the Sign of  $\triangle$ t

abs.db.	<b>≤</b> 250	>250 to ≤ 500	>500 to \$1000	>1000 to \$ 5000	> 5000	Total
≤ 3	1	2	3	4	0	10
>3 to ≤ 6	1	1	3	7	0	12
>6 to £10	0	0	0	3	· O	3
<b>&gt;</b> 10	0	0	1	4	8	13
TOTAL	2	3	7	18	8	38

TABLE 2.3 Relation Between the Intensity of the PCA Events and Peak Intensity of RF Emissions

Table 2.3 does not indicate any good correlation between the intensity of the PCA absorption and the peak flux of the RF emission.

#### 3.2 FLARES FOLLOWED BY VERY SMALL PCA EVENTS

We have ten events in this group of small or very small PCA events including seven events derived by Gregory based on the use of high sensitivity vertical incidence back scatter sounding of the lower ionosphere at a frequency of 2.3 Mc/s.

The times of both flare maximum and RF peak flux at one or more of the frequencies in the range from 1500 to 3750 Mc/s are known for nine of the events. In this case we find only two events where the RF maximum follows the flare maximum ( $\Delta t \ge 0$ ) while major RF bursts were reported (flux  $\ge 500 \times 10^{-22} \text{ W m}^{-2}(\text{c/s})^{-1}$ ) for three of the cases where  $\Delta t$  was negative.

Flux	Unknown	Positive	Zero	Negative	Total
∠ 500	1	0	0	4	5
> 500 to < 1000 > 1000	0	1	0	2 1	3 2
TOTAL	1	2	0	7	10

TABLE 3.1

\$\triangle t\$ for Flares Followed by Small PCA Events

Because of the probable very low absorption during these events we will combine these flares with those not followed by PCA events.

#### 3.3 FLARES OF IMPORTANCE 3 AND 3+ NOT FOLLOWED BY PCA EVENTS

Thirty-eight of the flares of importance 3 or 3+ were followed by important PCA events and ten more have been assoicated with very small events. In addition three of the PCA events occurred within three hours of the associated PCA flare and may have contributed to the intensity of the PCA event (flares 95, 100, and 137) although in two cases (95 and 100), the start of the flare followed the reported start time of the PCA. This leaves 91 importance 3 and 3+ flares that were not followed by a known PCA event (Table 4).

We find that eleven flares occurred during the normal observing times of two or more RF frequencies and eight during the normal observing time of at least one RF frequency with no increase in flux reported.

There was no known RF observations at the times of five flares (numbers 7, 11, 26, 28, and 32).

The times of flare maximum and/or RF peak intensity is not reported for 13 flares. The values of  $\Delta$ t for the 54 flares for which the times of flare maximum and RF peak intensity is negative for 49 of the flares. If we assume that at least one observatory was observing at the time of the 11 flares that occurred during the observing times of two or more frequencies, we have 60 (and possible 68) cases where  $\Delta$ t was negative or no RF flux was emitted as shown in Table 4.1. We see that for the importance 3 and 3+ flares that were not followed by a PCA event there was probably no RF emission from eleven (and possibly 19), and the RF emission was less than 500 x  $10^{-22}$  W ( $m^2$ c/s) $^{-1}$  for 47 of the flares.

Peak t Flux	No Flux	Unknown	Positiv <del>e</del>	Zero	Negative	Total
No known cm obs.		5				5
No flux reported	19					19
<b>5</b> 500		10	4	. 1	32	47
> 500 to		2	0	0	5	7
≥ 1000		1	0	o	12	13
TOTAL	19	18	4	1	49	91

TABLE 4.1
Relation Between the Intensity of the RF Flux and  $\triangle$  t

We also find that 4 was negative (the time of the RF flux maximum preceded the time of flare maximum) in 92% of the cases where the times of both maxima were known.

The four cases (Flares 34, 48, 54, 58) where  $\triangle t > 0$  and the one case of  $\triangle t = 0$  will be examined in detail later.

If we include the nine cases where (Table 3 and 3.1) very small or doubtful PCA events were reported with the 65 cases from Table 4 and 4.1, we find that in 90% of the cases the time of RF emission peak precedes the time of flare maximum.

PROBABILITY THAT AN IMPORTANCE 3 OR 3+ FLARE WILL OR WILL NOT BE FOLLOWED BY A PCA EVENT

Table 1.2 shows that if we include all cases where  $\Delta t \geqslant 0$ , the flare was followed by a PCA event in 79% of the cases. Similarly, when  $\Delta t < 0$ , or no flux is reported when the flare occurs during the normal operating time of two or more frequencies (11 flares), the flare was not followed by a PCA event in 90% of the cases.

Flare & t Type	Total Number	Unknown	Posit:	ive Zero	Nega		Flux ported
Non-PCA Flares	91	18	4	1	49	11	8
Very Small PCA	10	1	2	0	7	0	
Important PCA	38	4	23	3	8	0	
Others * Three	3* of the Po	CA events	may have	been influ	enced by	two major	flares.
TOTAL	142	23	29	4,	64	11	, 8
ACCURACY PROBABILITY				79%		90%	

TABLE 1.2 Classification of all Importance 3 and 3+ Flares in Terms of the Algebraic Sign of At

#### 4.0 SUNSPOT GROUP AREA AND MAGNETIC CLASSIFICATION ON FLARE DAY

#### 4.1 SUNSPOT AREA ON FLARE DAY

The 142 flares of importance 3 or 3+ were observed in 101 different sunspot groups. Of the 109 flares reported during 1955 through 1959, 74 occurred in sunspot groups that on flare day had an area equal to or greater than 500 millionths of the solar hemisphere. The distribution with sunspot area is shown in Table 1.3.

Twenty-three of the 25 major flares that were followed by a PCA event occurred in a sunspot with an area greater than 500 millionths.

Sunspot Area	<b>∢</b> 500	500 to	1000 to <1500	1500 to <2000	\$2000	Unknown EL, or Not Observed	WL
Number of Flares	30	33	22	14	5	5	
PCA Flares	2	12	7	3	1	0	

TABLE 1.3
Distribution of Sunspot Areas for All Importance 3 and 3+ Flares and Those Followed by a PCA Event

During this period 23 of the sunspot groups were the source of two or more major flares. In eleven cases at least one of the flares was followed by a PCA event.

We have limited the correlation of flares to sunspot area to the period through 1959 for which we have Greenwich Observatory daily sunspot area data.

Year	< 500	500 to	1000 to <1500	1500 to <b>4</b> 2000	<b>≥</b> 2000	Unknown
1955	0	3	1	0	0	0
1956	6	5	14	5	0	1
1957	11*	10	4	2	3	2
1958	4	9	3	4	0	0
1959	9	6	10	3	2	2
1960 <del>**</del>						
1961 <del>***</del>	2	1	1	1	0	6
1962***	1	0	1	0	0	0
1963 <b>***</b>	0	1	0	0	0	1
TOTAL	33 <b>*</b>	35	24	15	5	12
PCA	3	14	8	4	1	1

<sup>\*</sup> One flare (No. 57, Table 1) occurred in a region with no known sunspot groups.

TABLE 1.4
Distribution by Years of Importance 3 and 3+ Flares with the Area of the Sunspot Group on Flare Day

If we use the U. S. Naval Observatory area data for the years 1961 through 1963, we find that 79 of the 112 flares (approximately

<sup>\*\*</sup> Flare day sunspot area data not available for 1960.

<sup>\*\*\*</sup> Flare day sunspot area data after July 1, 1961, taken from U. S. Naval Observatory daily reports.

71%) for which the sunspot area is known on flare day have an area > 500 millionths of the solar hemisphere. The distribution of the importance 3 and 3+ flares that were followed by a PCA event is given in Table 1.5.

Year	<b>≰</b> 500	500 to <b>≼100</b> 0	1000 to <1500	1500 to <2000	<b>≯</b> 2000
1955	No PCA	Events			
1956	0	2	0	0	0
1957	2	6	1	1	1
1958	0	3	1	2	0
1959	0	1	4	1	0
1960*					
1961**	ı	, <b>1</b>	1	ı	0
1962	No PCA	Events			
1963	0	1	0	0	0
TOTAL	3	14	7	5	1

<sup>\*</sup> Daily sunspot areas not available.

TABLE 1.5
Distribution of Importance 3 and 3+ Flares Followed by a PCA Event with the Area of the Sunspot Group on Flare Day

<sup>\*\*</sup> Flare day sunspot area data after July 1, 1961, is taken from U.S. Naval Observatory daily reports.

#### 4.2 SUNSPOT GROUP MAGNETIC CLASSIFICATION ON FLARE DAY

The magnetic classification on flare day of the sunspot groups associated with the importance 3 and 3+ flares is known for 121 of the 142 flares as shown in Table 1.6.

Twenty-seven of the 30 PCA flare sunspot groups for which an area on flare day is known had an area greater than 500 millionths of the solar hemisphere. Both the sunspot magnetic classification and the sunspot area are known for 27 of the PCA flares. Nineteen of these had an area greater than 500 millionths and a 7 or 37 magnetic classification.

Both the flare day magnetic classification and area are known for 93 of the importance 3 and 3+ flares. Forty-three of these had a for fr magnetic classification and an area greater than 500 millionths. Only 44 per cent of the importance 3 and 3+ flares from large and magnetically complex sunspot groups were followed by PCA events.

	. 4	4 4				Unknown	
	8	BY	B	×	WL	EL	Not Rep.
1955	ı	0	3	0	0	o	0
1956	5	2	10	2	2	0	0
1957	2	11	6	3	2	1	7 <b>*</b>
1958	1	4	14	1	0	0	0
1959	11	7	8	1	0	2	3
1960	5	3	4	5	ı	0	ı
1961	0	3	3	2	2	0	0
1962	0	1	1	0	0	0	0
1963	0	1	1	0	0	0	0
TOTAL	25	32	50	14	7	3	11

<sup>\*</sup> One flare (No. 57, Table 1) occurred in a region with no known sunspot group.

TABLE 1.6
Distribution of Importance 3 and 3+ Flares
with Sunspot Magnetic Classification on Flare Day

	Y	ߥ	ß	ø	WL	EL	Not Reported
1955		None					
1956	1	0	0	0	0	0	ı
1957	2	5	ı	1	0	0	2
1958	0	2	14	0	0	. 0	0
1959	4	ı	1	0	0	0	o
1960	2	1	1	1	1	0	1
1961	0	3	ı	0	1	0	0
1962	None						
1963	0	1	0	0	0	0	0
TOTAL	9	13	8	2	2	0	3
TOTAL, All Major Flares	25	32	50	14	7	3	11
Percent with PCA	36	41	16	14	28	0	27
Small PCA (Table 3)	2	3	2	2	0	0	1
All PCA	11	16	10	4	2	0	4

TABLE 1.7
Distribution of Importance 3 and 3+ Flares
Followed by an Important PCA Event with Sunspot Magnetic Class

The sunspot magnetic classification on flare day is known for 121 of the importance 3 and 3+ flares. Fifty-seven (47%) were associated with sunspots that on flare day had a 7 or 88 magnetic classification. Only 22 (39%) were followed by a PCA event.

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	The flare	on 7/20,	/61 at 15	24 UT, Imp	The flare on 7/20/61 at 1524 UT, Importance 2, is considered by Lainbach to	1s cons	dered by	Leimbach	\$		

The flare on 7/20/5 at 1284 Up, importance 2, is considered by Latinach to be that source of PKA 55 although the importance 3 flares at 1633 and 1826 may have contributed to the proton stream.
 Plares with Importance 3+.
 Plares with Importance 3+.<

Zurich	Class	<b>Æ</b>	Ê	*	M	<b>143</b>	<b>6</b> 1	443	t	5	•	=	æ	=	=	<b>_</b>	==	×		ы		۴.		×	_	<u>.</u>	٩	,	ь	,	۵.		U	U	Δ	=	۵	-,	<u>.</u>	ь.	54.	ນ		۵	=	5	se se	
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Sunapot	Class	Ø	<b>)</b> Co	Ξ	(88)	E E	( <u>6</u>	98	3	} {	$\hat{\epsilon}$	ઉ	(£	(R)	$\mathfrak{S}$	ı	$\mathfrak{L}$	(£)	Ē	E	ş	(DD)		<b>}</b> - '	<b>&gt;</b> 0	<i>حد</i>	ď	West Li	g	8	9 صر	C;	×0 %	βP	$\alpha b$	٠ صد	Q.	$\sigma' \rho$	B	, ;	g,	B		g g	Эb	rs	14	
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	Date	4/05	80/1	4√13	60/5	o1/5 <b>⋖</b>	€5/10	2/11	9179	or/o	₽6/18	P7/10	-1/10	7/13	41/14	11/2	91//4	<b>1</b> //16	1/21	8/13	92/11▲	11/30	961	n/1	2/55	10/4	<b>4/28</b>	5/OF	90/9	60/5	5/13	10/9	6/52	92/94	92/9	12/9	10/29	11/06	o1/114	21/11	11/15	15/02	1961	3/56	92/4	<b>t</b> c/: <b>▲</b>	11//1	
Serial	No.	8	76	84	88	***	8	*	. 8	÷.	* &	*66	100	101	102	103	101	105	106	101	108	100		91	Ħ	112	113	777	115	911	11,	118*	119	120	121	122	123	154	125	126*	127	128		123	130	131	132	1

	Also a very great buret at 200 fe/s,) 2:000 peak.	Also a very great buret at 200 Mn/s, 22500 peak.	to cinterwations at other on, wavelengths.	Also 9500 Mafe start ORM lasting 877 mis. 630 mosth	flux. Occurred increase, not reported in 185 at 2950.	Also becase at 9900, 9905, and 1000 Me/s and at all moles weekingths.		This can'll This was included in the beside 11ft.  The act desired by Addresses of balloom allitude and by man of the determines of the Immemberic scular  bedeation.		the flare ensentable vith this small POA reported by many chaerese seemered in the same region as the flare 46, Puble 3.	is the instrume reported at my wavingth associated at the tar face are at 170 th and the face at a 170 th and the face are at 170 th and a section of the face are at 150 th associated with the face are at	Major bursts as all mater versionstabs. Also 9400 Mc/s with & b = 01. MET chestrad increase (not reported 1M).	Oit, was the only observatory at an werelangth observing at the time of the flare. Large bursts at moter werelangths remoted peak intensity with Q.t. = '95, and '120.		impler bursts at all mater verslangths with $\Delta^{\pm}$ ranging from 0 to 47.0. This flare was followed only 30 minutes after flare		Also 9000, 1000, and 1AgO 90/s.	Also major bursts at 9000, 9900, 1800, and 1000 Mc/s all with $\Delta t = 0$ , or +2. A great burst at 200 Mc/s.	Also plots with Alt = +1. Major bursts at metar weetlangth: Med Data not published in 1Mf. MRI	Also 1450 Me/s, flux 206 At = +15, and 0400, 9500, and 1000 Me/s.
1	Ž.		ř			\$\$ <b>22</b> \$	.a	×8.	£.	171	₹.	₹ ¥	÷	\$		£	000	۰۰۰	£4.	115
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įs	:	3	å			0440	1737	<b>§</b>	1316	0300	22	1335	1998	1642	1003	0110	1000	9	1450	1200
Durekion P Min.	8	ź	ğ			£	193	<b>K</b>	<b>6</b> 8	*		8	218	6	280		\$	239	981	æ
Start Duration UT Min.	<b>a</b> go	*	9			27.00	21/1	a	1297	8	1656 1608 2608	व्या स	7367	77.97	7460	230	6636	8	म्ब	800
Pate	}	8/31	ISS.	10/4	Ĉ.	1/03	<b>8</b>	8/26	θ/π	17/6	81/6	17/6	92/6		3/23	1/01	<b>1/3</b>	9/16	22/9	92/8
Flare Serial	L	=				<u>*</u>	3	3	1		<u>\$</u>	<u> </u>	2	<b>\$</b>	š	*19	38	•17	Ę.	<u></u>

TABLE 3
DEPORTANCE 3 AND 3+ PLARES WITH REPORTED SMALL POA EVENTS

									rage	NO. 25	
POTES AND REMARKS		The flare reported by Sec Peak only, assigned importance 3+ by Wolseth. This small PCA reported by Be, DLP, DLP, DLP, and S, based on ionospheric fain data. The RP burst is a small but long enduring rise and fall in flux.	Also at 9400 Mc/s, flux = 663, £t = -4. PCA reported by Juc man at 78 mly. Another at 059; Another imp 3 fluxe (No. 80, Table 3) in the same region	For reported by B1, Q, with start times 11/2200 and 12/0700 respectively, such by JC and 8 with 13/1600 and 12/0700 respectively, such by JC and 8 with 13/1600 and 13/1600 and times all of importance 1 reported between 11/2040 and 13/1201.  None with important radio bursts.	This very small PCA reported by Gregory only. No spectral observations at the time of the flare.	Also a major burst at 9100 Mc/s, leating 63 minutes, peak flux > 3400. Wary small PCA reported by dregory only. No spectral observations at the time of the flare.	Small PCA reported by Gregory only. HBH observing no report in IAU Bulletin.	Shall FCA reported by Oregory only. Nag. observing. No resport. The type in reported by Ward it requency 33 to 15 No. 4s. reported unclassailed by Ft. Davis and thinks and Maracick continus the Type II." Hothigan and Maracick continus the Type IV for 119 minutes. This was the third importance 3 fare (flares 119, 120 & 122) from this region. Flare 120 (Table 4) did not produce a FCA event. The redio burnts were moderate at on FCA event. The radio burnts were moderate at on GA event. The radio burnts were nowlength, starting almost simultaneously with the flare but of whort duration and negative & t.	Smill PCA reported by Gregory. No spectral observations at the time of the flare. A probable Type IV based on the large bursts at all frequencies from 23 to $9400~{\rm Me/s}$ .	This burst had 3 major peaks. Also major bursts at 9400 Mc/s. No spectral observations at the time of the first of a series of major solar events from Plage 5925 witch was the source of three major PCA wents (Members 40, 50 and 51) and four small events responded by Gregory only (034, and 036), between 11/10 and 11/21 when the region was about 20 degrees beyond the west limb.	Inis small PCA reported by Gregory only is in a region which is a return of the very active region 5425.
I	*	-15	-1.5	-18	**	-13.5	6.6.0	-2.0	+25.6 +19.0 +15.0	-2.4 +62.4 +77.6 -1.3	-0.5
	Energy			13			91	71		•	23
	Flux	(61)	\$ 98.5 \$	220	<b>&gt;</b> 20	3100	610	140	583 1000 <b>&gt;</b> 800	365 100 360	330
DICTES	Min.	240.0	5.0 12.5 189.0	<b>X</b> 35	200	88	241 88	38	1055.2 273.5 1049 50 1045.0 \$235	112	27
SINGLE PREGUENCIES	ž.	1915	0858.5	2108	; ;	0846.5 ~300	1207.0 1209	2158	1026.5 1055.2 273. 1036 1049 50 1026.0 1045.0 ≯235	1120.6 1135.6 1135.8 1021.7 1022 >100 1131	1837.5
8,000	get E	1750	0855 0855 0852	5056	0700 0710	0822.4	1148.5	2340		1016.4	1828 1855
	2	2800 80	1500 2980 3000	2800 s	2000 S 3750 S	1500 C+ 2980 C+ 3000	1500 2800 7980 3000	2800 8 3750	1500 C+ 2980 B 3000 C+	2980 C	2900
BPECTRAL	Dur/Imp.		ı	2105 170/3	ı	(0837)	1215	2150 44/3	See Note		1833.5
92	Imp. II	1605 2008 3/2	<u> </u>	2- 2103.3	:	:		215/3		N.	3 1834 16/3
VE PADE	Min. D		50 5	24 2	88	75 3	67 2	138	£	8	100
SHORTWAYE PADE	Onset		0855	2100	940	0837	1203	2141	1029	1022	1830
×	Intensity	æ	so	\$2	82	. <b>8</b> 8	8	EX.	8	EA	Ř
BSORPTTO	Duration Hrs.	*	:	<b>%</b>	57.	144	84 4	78.	192	₹ <b>.</b>	144
POLAR CAP ABSORPTION	Serial Start No. UT	1800	15%	2200	88	1400	1733	2300	1200	1800	3 05X
8	Max. Serie	83	816	819	016	, to 020	955	653	430	1 031	3 038
	No. Rep.	1/1	1/6	2/5	1/1	01/81	4/8	2/11	z/n	1/6	11/3
	Position	N20 W30	N16 W61	N22 B03	811 252	N29 E46	N21 E06	NET WE?	N22 826	N28 529	N26 874
FLARE DATA	ž.	1930	0060	5715	0734	0060	1215	2156	1030	1023	1838
-	Start Duration Ur Min.	360	108	195	197	954	239	125	185	231	325
	Start	1605	2480	2040	4010	<b>128</b> 0	1131	2140	9 1026	0 1009	5 1825
	Date	7/2	1/26	77/T	2/00	10/9	6/25	6/27	10/29	11/10	12/05
	are rial lo.	33*	79	011	116	118	611	753	123	125	128*

TABLE 4

DEFORMANCE 3 AND 3+ PLANES

VITE NO ENGINE ASSOCIATION PCA EVEN

	Τ	-	FAR DE	Α			<b>30</b>	WIE F	426	Start UI			DENIA 1		76 G	26. 21	n pundit		
Flare Serial		Start	Duration	Max.		No. /No.	1	har.		Der / Sep.	1			-	Der.	-			
No.	Date	ÚT	Ma.	UI	Position	No/ im.	Connect	Hin.	1	- 1 N	-	Ž-	UT	ŪŢ.	His.		learny.	₽ŧ	Rotes
	1955						1			İ	İ								
1	1-16	2130	50		#33 <b>#41</b>	1/1					2800	<b>C</b>	21.05		30.0	(>133)		M <sub>12</sub>	
2	6-18	1218	57	1232	<b>323 1125</b>	4/2	1227	₩8	1		2800	<b>C</b>	1222		53-5	(> 5%)		M2	1
3	11-12	1116	43	1133	<b>E27 E37</b>	2/1	1327	23	2+		2800								Probably observing. Hornal observing time with quarter 1955; 10-2400 UP
4	12-03	1058	93	1112	M22 M10	2/1	1105	20	3		2800								
_	1956					•													
5	1-19	0535	100		E22 E19	1/1	0558	32	1+		3750								Horanl observing time 00-0600 UT
6	2-14	0538	112	0557	M21 #33	3/2	0532	116	3	0554	3000	œ	0541						Time of peak flux not reported.
										26/-	3750	8	0541	0553	75.0	2720	66	-4	Also 0538-0538.7, SD, Peak Flux (15
7	2-17	1100	102	1750	20 HOF	8/3	1105	**	3										None observed at on wave length.
9	2-29	2220	49		\$30 ¥21	2/2	2228	72	3-		2600	80	2217		23.0	(525)		M12	Times of flare max. & MF peak not reported.
10	3-02	1220	80		M21 W64	1/1	1158	34	s		2800								Hormal observing time 1st quarter
						- 4-													1956 12-2300 199
12	4-09	0940	70 94	1000	\$22 \$27 \$24 \$18	5/2	0945	70	2	<b>4</b>		_							Hone observing at ca wavelengths
13	5-17	2230	76	2305		1/1			_	(2234)	2600	500	2230	2252	50.0	(35)		-13	
15	8-29	<u>0937</u>	76	0956	#30 W63	8/2	0936	æ	2		1500 2600								MII may have observed. Survive, normal observing time
15	9-05	1445	22		825 E77	1/1													10-2300 UT
16	9-14	0813	<i>42</i> 54		S22 129	•					1500 2800								So MF reported at on wavelength.
17	10-01	0756	60		545 MB	3/1					1500								So MF reported at on wavelengths.
					-	1/1					1500								No MF reported at any frequency.
16	10-07	0400	62		9624 BO7	1/1					3000	0	0411	0413.		273		*1	Time of flare max. not reported.
	l										3790	66	0446 0350	0447	6.0 2.3	363 (32)		H <sub>12</sub>	
19	10-11	0955	78	1026	M22 W56	5/2	10375	50	3-		1500							_	
20	<b>~11-0</b> 7	1109	174	1135	S17 B32	6/20	1327	153	3-	(1103)	1500		1106		113.0	(465)		M <sub>2</sub>	Time of peak intensity not reported
21	11-14	1037	230	1055	820 W55	3/2	1.037	76	2+	(1035)	1500	<b>a</b>	1032		73.0	(820)		Mo	Also 9400 Mc/s starting 1035 DT lasting 115 mins., peak smoothed
																			lasting 115 mins., peak smoothed flux 1045.
22	77-50	1002	188	1020	\$15 <b>¥5</b> 6	6/2	1007	59	3-	(1909)	1500	9	1010		215.0	(>1500)		M <sub>2</sub>	Also 9800 Nc/s start 1000 UT last: 202 mins., peak assorbed flux>500
																			Major bursts at meter savelength.
23	12-06	1405	9		S21 B41	1/1	1338	38	3-		5800	80	1336 1342	1339.	5 6.0	127 10		*1	
24	12-17	1535	90	1551	S24 #52	1/1	1554	75	3		2800	530		5 1545.	95.0			-5.5	
"	12-17	±232	90	2771	324 #32	1/1	1774	13	3		2000	24)	1555	2 1242.	100.0	50		-5.5	
25	12-22	0955	63		#32 #90	1/1					1500								No bursts reported at any waveleng Normal observing times 07-1400 UP
																			st 1500 Mc/s
	1957																		
26	1-06	1038	205	1126	S21 🗯	6/1													No observations at on wavelength.
28	1-23	5310	48	2314	<b>#17 <b>#2</b>7</b>	1/10													No observations at on wavelength.
29	1-24	1225	89	1241	N16 W32	4/1	1235	35	2		2800	<b>SD</b>	C1235		<b>&gt;</b> 10.0	(>120)			Also moderate bursts at 200 & 536
30	1-25	0520	17	0526	520 <b>45</b> 0	16	~~~		1										Mc/s.
,		0,20	•	0,20		1/1	0528	20	1		3000 3750								Moreal observing time lat quarter 00-0600 UT at 3000 and 3750 Mc/s.
31	1-31	0358	112	0436	<b>30</b> 5	1/10	0356	84	1	akat 17/3	3000 3750	æ	0400	0435	>120.0	234		-1	
, 2	2-08	0832	43	0836	MA Inc	1/1				11/3	3150								So flux reported.
_		عرت		w.y.	— w	1/1													No observations at on wavelength, small bursts at meter wavelength.
34	2-26	0005	255	001#	<b>135</b>	1/1	0050	110	1+	0009 0029 17/3+ 57/-	3000	<b>a</b>	0000	00 <b>\</b> 5	>50.0	224		+31	Flare was reported by Mitaha only
35	<b>4</b> -02	1002	10		and war	1/1				75. 711-	1500	SD	1036			(129)			(IMU) No report at 2980.
37	-16	1040	140	1105	910 E05	10/2	1044	76	3		1500	æ	W50		2.0	(129)			!
·	-		240			+4/€		,00	,		2900 2980	8	1040 1037	1046	50.0	(1650) 1600	87	-19	HEI was observing at 1500 Mc/s but not included in IAU Bulletin
38	4-17	1006	72	1022	<b>300</b> 276	6/2	100%	79	3		1500	4	1037		50.0	THU			Plane commend destine normal at
	-		15			J,2	2004	19	3		5000 1200		1006					_	Flare occurred during normal ob- serving time at 1500, 2000, and
											2900	w	1000		4.0				3000 Mc/s, so flux reported. Sumrise at 2800.
39=	- 4-17	2000	180	<b>211</b> 6	<b>5</b> 0 <b>2</b> 69	1/1•	1937	163	3+	2032 2011	2800	G28	2006	2042	79.0	600o	5 <b>4</b> 6	-34	This flare reported by Honolulu
Į										44/2					. ,			,-	is not included in the IAU Bullets
							l												I

TABLE & (cont.)

											TABLE & (cont									
F	Ŧ			PLARE DAT	A			5 <b>808</b> 0	WATE !	PAIR	STACHAL TYPE Start. UZ		81	BULL PR	01.00		A STATE OF	1001		
Flare Serie	إنه		Stert	Duretion	Max. UT		No. No. No. Inc.	Onest	Per.	•	Sear./Seap. Rija		Type	Onset	Max. UT	Dur.	Peak Int.	Beerry	At	liotes
#0. #0	-	-ok	1134	<b>Rin.</b> 20	UI	Position NO.2 139	1/1		MID.	<b></b> -	14	1500 2000 2800 2980 3000				A124				This flare reported by Moscov only, occurred dering the normal 3rd quarter observing-time of at least five M observatories. Here re- ported a flax increase.
113	7-	-22	0953	117		#15 # <b>5</b> 1	1/1					1500 2000 2980 3000 3750								Fare reported by Moscow only. Many small bursts reported between 0946 and 1006 UT at mater several the sormal operating times of the listed frequencies, also at 7900. 3975, and 9800. No flux reported at may ins werelength.
45	9	<b>-0</b> 3	1412	195	1429	#83 M30	13/6	1420	103	3	(1417)	1500 2600 2980 3000	50 50	1420 1417	1425 1424	80.0 25.0	(509) (1350)	51	-4 -5	A large burst was also recorded at 9000 Mc/s _t w-6. Red recorded an increase at 2900 Moderate to large bursts were reported at most mater wavelengths.
₩	9	<b>-</b> 10	0223	37	0250	#14 #11	1/1					2000 3000 3750	888	0222 0223 0222	0226 0226 0226	35.0 35.0 35.0	(15) 349 (38)		-22 -22 -22	This flare was reported by Sydney only.
48	7	-18	1026	347	1045	H23 E30	12/3	1030	104	3		1500	8	1100	1703	7.0	(157)		+18	All of the RF Bursts associated with this flare started after the
												2600 2980 3000	8um 50 50	1102 1100	1705	2.5	308		+17	reported flare maximum at 1045 UT, even the there was a major SMF starting a few minutes after the flare, all of the MF bursts were small and of short duration.
50	٥	<del>-</del> 10	9350	125	0410	#23 B02	5/1	0359	54	3	ok27 183/3	2000 3000 3750	988	0402 0403 0359	0406 0406 0406	10.0	(254) 1410 (10 <b>8</b> 0)	12	4	
73	٥	<b>3</b> 0	1657	3 <b>3</b>	1706	. N25 ¥37	5/3	1700	40	3		1500 2000 2800 2900	200	1658 1658 1700 1705 .	1658.4 1710 1702 5 1706.3	70.0 6.0	(18) (30) (77) (120)		-7.6 +4 -4 +0.3	The Ef emissions were reported at 2800 Me/s only and consist of several small bursts super posed on a rise and Fall in Flms. No bursts were reported at the other on unvalength and none were reported at meter wavelength.
5 <b>4</b>		0-16	0152	10	C152	S25 <b>B</b> 21	1/1	<b>01</b> 50	20	2+		3000 2000 3750	CD #CTD	0151.0 0151.0	9 0152.8 9 0152.8	3 3.0 2 3.0	92 (165)		+0.8 +0.2	This flare was reported by Ritaka only. Small bursts were also reported at 9000 and 1000 Mc/s.
āĠ	12	1~24	0646	194	0911	514 E37	7/1	0907	32	3-	(0857)B	1500 2980 3000	999	0857 0859 <0006.5	0904 0906.5	40.0	(251) >998		-7 H2 -4.5	
5"	1:	1-2+	<u>9945</u>	315	<b>3213</b>	941 <b>5</b> 63	1/1•				0099 4/2	2000 3000 3750					-			This flare was reported by Sydney only. Bo RF bursts were reported at the selected Frequencies. A burst at 0500 Me/s lamted for 120 minutes with speak flux 400 and 01-5 was the only reported emission although Sydney reported a Type II which started before the flare many the flare many the flare many than
5 <b>8</b>		.958 -25	0915	115	1005	524 W69	17/1	0938	74	3		1500	ca.	0948	1015	51.0	122		+10	Also max at 1006, 1011 with at
	-											2980 3000	ED 80	0932 0949	1015	40.0 83.0	278 358		+70	Also max et 0957, 1002, 1005 with att: -8, -3, 0, resp.
59	3-	-01	0905	્રસ્	0917	\$11 <b>₩</b> 6	5/1	0913	12	24		1500 2980 3000	8D 8D	0911 0913.5	0914	21.0			•3 #2	MMI not reported in IAU
60	3.	-03	1005	165	1020	816 <b>26</b> 0	8/1	1070	95	3+	(1005)B	1500 2980 3000	SD	1005		28.0	1338 1015		<b>M</b> 2 -5	Great bursts at meter savelengths.  Max reported by Jod.
61	3.	-05	0500	92	9540	<b>\$12 \$46</b>	1/1					2000 3000 3750								This is a questionable importance 3 flare.
63	4-	-07	1010	125	1025	M15 M33	5/2	10076	69	2+		1500 2980 3000	CD SD	1013 1014	1016.5	36.0 8.0	564 736		-8.5 Ng	
÷4	5-	-01	2115	86	2130	819 <b>2</b> 15	4/1	2130	25	1+		2800	SD SA	2127.9 2127.9	2157 2128.5	5.0 5.0	(න) (1%)		+27 -1.5	This week gradual rise and full of long duration has the superposed minor burst with a at =-1.5. The large burst at 9500 (peak>820) Also has a at = -1.5.
65	5-	-05	0356	61	0415	518 W29	3/2	0407	53	3	(0412)	2000 3000 3750	899	0412 0412 0412	0414.9 0414.8 0414.1	12.0	(470) 1200 (900)		-0.1 -0.2 -0.9	Also max at Dklk.8, Okl9.6 with <u>At: -0.2, +4.6, resp.</u>
66	6-	-19	0940	150	1010	#14 W21	14/5	1005	25	2	·	1500 2800 2980 3000	C) S0	0943 0942.6	0944.5 3 0944.2	107.0 5.0	116 (14c)		-25. -25.	Small bursts lasting from 30 to 200 minutes reported at all meter wavelengths. HET and Hed. also observed increase, not reported.
69	a-	<b>⊢0</b> 4	0409	121	0435	#30 #31	3/1	0422	96	3	See Bote	2000 3000 3790	868	0427 0427 0428	0432 0431 0431	10.0 20.0			-3 -4 -4	Also bursts at 1000 and 9500.  Sydney reports possible Type II 0839-0853.

TABLE & (cont.)

			AR SATE					. WAVE	PATE .	Antibal tors		RTN	21 P.	Ti Gui			e e e e e e e e e e e e e e e e e e e		
Flare						/ No.				Start UZ Dur./Esp.									
Serial No.	Date	Start	Buration Nin.	Max. UT	Position	No. / Max. Rep Yep.	Omet	Nur.	<b>39</b> .	Min/		Type	UT UT	Wax.	Min.		Berry	<b>≜</b> Ł	Botes
70	8-07	1457	123	1508	816 E71	10/4	1500	105	3*		1500 2600	SSO CA	1500 1500	1503 1503	57.3 15.5	(95)		-5 -3	Flare in the same region as the FCA flare No. 26 on 8/16.
											2980 3000	CM	1516	1522	12.5	(18o)		+14	MII and Not observed increase of 1500 and 1516 UT
74	9-18	0728	130	0830	S12 W53	8/3					1500	80	0826	0829.5	5.0	222		-0.5	
											3000	80	0827	0830.0	11.3	261		0	Also max. at 0835 UT
75	11-14	0036	ðΤ	0046	S19 <b>E</b> 51	2/1	0039	61	3		2000 3000 3750	8	0033	0099		(32)		-7 -6	Also small bursts at 1000 and 9400 Nc/s moderate burst at 9500 Mc/s.
70	11-24	1607	180	1621	S12 W08	3/2	1615	80	3-	1607	2800	BA	1613	1620		(285)		-1	Also at 1600 UT lasting > 30 minutes
77	12-11	1656	67	1703	518 <b>#5</b> 4	2/2	1700	36	2+	12/3 1705/ 1700	3000	æ	1658	1702	18.0	(340)		-1	
										36/2+ 12/2	3000								Summert at Jod.
76	1959 1-21	1700	50	1709	W20 X48	2/1	1702	41	2+	1718	1500								
										25/3	2600 3000	20) 200	1702	1708	26.5	600	12	-7	
èО	1-26	1027	158	1050	N16 W61	8/2					1500	9	1025	1037		765		-13	The small FCA (S16) was associated
											2980 3000	æ	1032		34.0	1099		N/Q	with the importance 3 flare at 0842 in the same region.
ð1	5-01	0352	56	0423	N12 1283	1/1	0422	25	3-		2000 3000	co _	0408	0422		(270)		-1	Also bursts at 1000, 9400 and 9500 Mc/s. Tok.not observing at 3000
82	2-02	1015	132	1039	NG2 N'30	7/3	1032	18	1		3750 1500	æ	0408	0422	17.0	(550)	6	-1	Mc/s Feb. 01 through Feb. 05. No NF activity reported between 0635 and 1817 UT.
						.,,					2800 3000								0635 mmd 1817 UT.
83	2-02	<u> 1816</u>	123	1824	1009 \$560	2/2	1817	85	3	1853 10/3	(2800	CD	1817	1821.2	65.0	30		-2.8	The R emission consisted of com- plex bursts separated on a long
											3200		1819. 1814.	7 1822.2 6	70.0			-1.8	but small rise and Fall in flux. Other 2800 Mc/s peaks at 1851.3 and 1910.3
95	2-18	1005	15		E21 E76	1/1					1500								It is doubtful if either flare
56	2-16	1142	96		<b>#21 E76</b>	1/1					2800 3000								85 or 86 has importance 3. They are included in Tuble 1 but not in Tuble VIII, Solar Activity Cat.
																			They are included in the McMath Working List. Both flares were reported by Kenzelhobe.
																			reported by management.
87	2-19	2032	132	2037	#17 A10	3/1	2030	40	1•	(2030)	2800 3200	80	2030 2030	2034.5	80.0 12.0	(200) 417		-2.5	The time of max, is indeterminate; however, times of peak flux at 3200
88	3-24	095â	207	1015	N29 W77	16/7	1002	107	3	(1003)	1500	æ	1005	1013	30.0	875		-2	and 9530 Mc/s indicate a magative & t
		<u> </u>		,	,	,			•	(===5,	2980 3000	<b>S</b> D	1003 1003		25.0 57.0	>1729		-3	BRI reports a max at 1012.
69	3-28	2113	122	2126	#24 H33	5/3	2121	24	1+		2800	SD	2121	21.24	7.0	(100)		-5	
90	4-05	2316	123	2327	W16 W67	4/2*	2317	93	3+		2000 2000 3750	ED .	2320 sunse 2318		9.0 -8.0	(580) (2300)	37	-4	Also major burst at 9500 Mc/s (peak flux 2960) and small burst at 1000 Mc/s (smooth peak 38)
91	<b>P4-</b> 08	0903	67	0921	R27 185	9/3	0913	32	2+		1500	SD	0915	0916	15.2	> 925	٠,	-5	,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,
											2980 3000	SD CD	0915 0914	0916	22.0 18.5	> 2150 1484		<b>H</b> 2 -5	Also observed by Gor. and Sim.
92°	4-13	0823	73	0840	<b>#27 £19</b>	16/2	0835	25	2		1500 2980	SD	0832 0832	0840 8040	12.0	232 282		M <sub>2</sub>	
93	<b>-</b> 5-09	0153	49	0150	M20 E78	1/1					3000			ter burs				9	Flares 93 and 96 occur in the same
		-	-			,-					2980 3000 3750	dur:	ing the	he flare normal frequenc	observ	red ing tim	45		region as PCA (34) associated with flares 94 and 95 on May 10 that were also responsible for the sea level
<b>9</b> 6	5-11	2005	104	2026	N10 E41	4/1	2015	67	3-	2020 2028	2800			2022		(900)		-6	event on May 11 at 0030 UT. Flare 96 which has all of the characteristics
										19/3+ 18/3									of a FCA flare occurred while FCA 34 was still increasing. However, Leinbach found no evidence that this
		2618	,	~	<b></b>	- 10				(0624)	1500								Flare did contribute to the PCA.  Also a major burst at 9400 Mc/s
97	-6-1c	3678	112	0628	<b>11</b> 5 £15	9/5	0623	*	2	(0624)	2000	SD	0623	0625.9	5.0	(1150	)	-2.1	(amouthed peak 2300 derations very short). He bursts reported at the
											3000 3750	œ	0623	0625.6	6.0	(1100	) 16	-2.4	other on wavelengths although the
98	-0-1ô	1134	86	1148	#16 W12	7/1	1138	22	2+	(1138)	1500	550	1138	1142		1025		-6	Also smjor bursts at 9375 and 19000
						:					2600 2980 3000	8	1139	1140	180	(1225	) 14	-8 -8	Mc/s with At= -9 Also Gor. and Sim.
101	<b>-</b> 7-13	0255	190	0410	<b>#1</b> 5 <b>#18</b>	1/1					2000	69	0249	0254	12.0	(24	)	76	Flares 101 and 104 occurred in the
											3000 3750	80	0249 0248.	0250 8 0249.5	3.0	246 (38	1)	-80 -80.	stare region as PCA's 35, 36, and 37, flares 99-100, 102, and 105.
103	7-14	1400	210	1452	S25 E37	14/1	1355	105	2+		1500 2800	M SD	1443 1443	1447 1446	32.0 30	340 (85	)	-5 -6	Channel by Red. Com and allow
											2980 3000								Observed by Med, flux and time not reported
104	7-16	1525	110	1616	¥14 ¥27	8/1	1610	28	2-	1616 7/13	1500 2800	SA	1613	1615	9.0	(350	)	-1	Observed by MMI, flux and time not reported
							Ì				2980 3000	n	2013	2027	7.0	(3)(	•	~1	Observed by Med, flux and time not reported
	l				_		I			l	1								1

TABLE & (cont.)

		1	LARE DATA				$\neg$	SHORT WAVE PADE   SPECTBAL TYPE					SINGLE PROGRACIES CONTROLLS BURSTS							T
	<del>                                     </del>							anux	MATE	PAUL		Art UT	<b></b>	3.0	OLE PIGE	LIDIC D	S CENTR		<del> </del>	
Flare Serial			Duration	Max.		No.	No.		_		3		1							
No.	Date	OT	Nin.	UT.	Position	No.	- I	Others	Dur.	ъ.	<del>- 17</del>	IN. IV	1.	Typ	Onest IV	Max. UZ	Mar.	Peak Int. Emergy	A+	Notes
106	7-27	2050	120	2115	E27 E26	6/2														<del></del>
				,		•		2105	23		2118 8,2		2800	SA SA	<b>2100</b> 2107	2111 2111	10.0	{ 20} 75}	+13 -4	A burst superposed on a long rise and fall in flux neither of which is very strone.
108	11-26	2006	84	2018	N12 E31	3/2	- [		35	2	2017 28/3	(2027)	2800	SĐ	5070		15.0	>225	W2	Flares 105 and 109 occurred in a region which during disk passage
109	<u>_11-30</u>	1720	106	1744	<b>307 206</b>	5/4		1735	47	3-	1741 29/3+	1739 351/3	2800	SD	1737.6 1745.6		18.0 60.0	(9175) 20	<b>₩</b> 2	produced four importance 2º Plaves and one importance 2 Plaves that were fallowed by Type II and/or Type IV spectral missacions. The importance statement of the product of
	1960																			
זוו	2-22	1352	68	1400	MOS 3941	4/1		1358	<b>F</b> 2	3-	1358 13/3	<1356 >4/2	1500 2800	C			\$ 50.0 28.0	319 340	→.6 -1.0	Also 9400 Mc/s 1353.8 UT, 59.8 min. Peak flux 455, _at = -1.4
120	6-26	0428	57	0436	M20 W08	2/2		0 <b>43</b> 2	56	1+			2000 3000 3750	e c			10.0 7 8.0 2 10.0	92 1422 225	-3.8 -3.8 -3.8	
121	6-26	2358	59	2415	S08 #34	5/2		2403	67	2-	2404 4.5/3	2413 36/3	2000 2800 3750	<b>s</b>	2402 2405	2412.	>28.0 5 50.0	13 50	<b>H</b> 3.5	Also Mag. In summet oscillation
124	11-06	€1752	158	1841	#13 B07	1/1	- 1	1708	67	1	1840		2800		1835	1836	25.0	28	-3	Flare reported by Honolulu only.
	1961										10/2		-							
129	3/26	1009	101	1035	515 <b>87</b> 4	16/6		1019	41	3			1500 2980 3000	c÷ c	1020 1026	1028.	7 110 14.0	9460 800	-6.3 #2	During normal observing time of
130	4-26	1646	179	1710	\$11 <b>2</b> 54	2/1		<b>165</b> 0	113	3			1500 2800		1640 1656.6	3 1702	237 8.0	32 18	<b>*</b> 3	HKI During normal observing time of HHI
131	S-04	2145	115	2213	S11 ¥56	4/2		2205	40	1•	2209 30.3/1	<b>2249</b> 72/1-	2000 2800 3750	c *	2205 2205 2207	2208 2208. 2209	8.0 8 9.0 6.0	110 95 80	-5 -4.2	Also 9400 and 1000 Nc/s
134	7-15	1433	296	1556	#13 E15	8/1		1512	113	,		1533	1500		1428.4	1426.	7 1.6	3	-92.1	Also <1435, max 1600.7, >240 min.
1							- 1					50/3	2800	e	1432	1436.	3 14.0	54	-79.7	flux 296, 4 t +4.7 Also 1432, max 1623, 458min.flux
							İ						2980	c	1433	1436.	5		-79.5	5h, 4 t 2+ 4,5
	1962						1												<del></del>	
139	3-22	2220	50	2241	R07 1236	1/1							2000	c		-	115.0			Sunrise   flare reported by Sec
1													2800 3750	-	2214		36.0 2 >15.0	35 29		Sunset records incomplete
140		1734	235	1806	MO9 E05	5/2	1	752	106		1844.4 B.6/2	1839 63/2-3	2900	:			294.0 5 11.0	25 20	-2.5	The gradual small rise which lasts for nearly six hours has the very small superposed burst with a &t = -2.5. No other single fre- quency at either or or meter save- length are reported.
141	9-16	0325	172	0422	<b>H11 B</b> 57	3/							2000	es	0419.5 0419.5		1.0	55	-2.4	Also 0400 Mc/s at 0419-0423, peak flux 110, 45 t = -2.3.

## APPENDIX A

For the sake of convenience we have included Tables 1, 2, 7, and 8 from Source 4, as Tables A-1, A-2, A-3, and A-4.

# TABLE A-1 PCA'S WITH ASSOCIATED FLARES AND SOURCES

No.	Date	PCA	Rise I	Duration hrs.	Abs.	Fla: Onset	re Max.	Imp.	Position	Δt	Range Start Times	Original Sources	Others
	1956	Deg.	True jii b.		40.		· ·	ш.р.	TOATCION		Marige Court Times	0000000	ocher 2
1.	Feb. 23	0400	18	123	13.0	0334		3	N23 W80	o <sup>h</sup> 23 <sup>m</sup>	23/0400-23/0600	B1, B2, Be, CJM, JP, S	M, Mo, WH
2.	Mar. 10	0900	36	160	3.5	0515		2	N16 E88	3h45m	10/0900-11/1600	B <sub>1</sub> , Be, CJM	ĸ
3.	Aug. 31	1430	14	69	4.9	1226	1246	3+(3)	N15 E15	2h04m	31/1430-31/1800	B <sub>1</sub> , B <sub>2</sub> , Be, CJM, JP, S	M, Mo. WH
4.	Nov. 13	2000	27	63	5.4	1430	1501	2	N16 W10	5 <sup>h</sup> 30 <sup>m</sup>	13/1400-14/1600	B <sub>1</sub> , B <sub>2</sub> , Be, CJM	M, Mo, WH
	1957										-,		
5.	Jan. 20	1500	16	86	4.1	1100	1119	3	\$30 W18	$^{7}\mu^{00}$	20/1500-21/1500	B <sub>1</sub> , B <sub>2</sub> , Be, CJM, DLF, S	BO, M, Mo, WH
6.	Apr. 3	1330	14	65	3.9	0825	0835	3	<b>S14 W</b> 60	5 <sup>h</sup> 05 <sup>m</sup>	02/2300-04/1200	B <sub>1</sub> , B <sub>2</sub> , Be, CJM, JP	BO, M, Mc, WH
7.	June 19	2215			Weak	1609	1613	2+(2)	N20 E45	6 <sup>h</sup> 06 <sup>m</sup>	19/2215-19/2300	Be, JP, <u>S</u>	во
8.	June 22	0500	44	115	5.0	<u>0236</u>		2	N23 E12	2h24m	22/0500-22/1000	B1, B2, K, RL, S	BO, M, WH
9.	July 3	0900	12	52	9.2	<b>[</b> 0712	0745	3+	N14 W40	1,542m	03/0815-03/1030	B <sub>1</sub> , B <sub>2</sub> , Be, Bo, DLP,	BO, M, Mo, WH
						<b>[083</b> 0	0840.	3+	N10 W42	0µ30m)		HG, HNS, H, JP, K, OH, PS, RL, S	1
10.	July 24	2015		27	2.0	£1712	1737	3	S24 W27	$3_{\mu}$ 0 $3_{m}$	24/1000-24/2400		BO, M, Mo, WH
,,	4	1600	10		2.1	11801	1828	•		chi.am	00 /2500 00 /0l/00	K, OH, PS, RL, S	no v
11.	Aug. 9	1600	10	50	3.1	0617 0810	0629	2	S09 E76	9 <sup>h</sup> 43 <sup>m</sup> 4h50 <sup>m</sup>	09/1500-09/2400 28/0400-28 <b>/&lt;</b> 2300	<u>B</u> 1, Be, H, JC, K, PS	BO, M
	-	1300	_		Weak		0955	3+(3)	S30 E35	_	, ,	A, Be, <u>Bo</u> , DLP, H, PS	BO, Mo
13.	Aug. 29	0000	7	27	3.2	28/2010	2024	3(2+)	S28 E30	3 <sup>r</sup> 50 <sup>m</sup>	29/0000-29/0500	A, B <sub>1</sub> , B <sub>2</sub> , HG, HNS, OH, PS, S	BO, M, Mo, WH
14.	Aug. 29	1300	12	58	8.2	1031	1052	3(2)	S25 E20	3 <sup>h</sup> 29 <sup>m</sup>	29/1300-29/1500		BO, M, WH
		23.25	30	1.6				>		- h - o#	(-) (	K, OH, RL, S	
15.	Aug. 31	1415	12	46	4.9	1257	1313	3+(3)	N25 W02	1 <sup>h</sup> 18 <sup>m</sup>	31/1415-31/1530	A, B <sub>1</sub> , B <sub>2</sub> , H, <u>K</u> , <u>S</u>	BO, M, Mo, WH
14	S M	1700	0	46	7.0	3 23 2		2(0.)	nal III	ahl-aff	m/1500 m/0100		
16.	Sept. 02	1700	9	410	7.2	1313		3(2+)	834 W36	3h47 <sup>m</sup>	02/1500+02/2100	B <sub>1</sub> , B <sub>2</sub> , B <sub>0</sub> , HNS, H, K, OH, RL, S	BO, M, Mc, WH
17.	Sept. 12	0500	13	57	0.5	11/0236	0300	3	N13 WO2	23 <sup>h</sup> 24 <sup>m</sup>	12/0200-12/2315	B2, Be, Bo, DLP, HG,	BO, M, Mo, WH
												HNS, H, JP, K, Kh, OH, PS, RL, S	
18.	Sept. 18	2000				18 [1658 1818	1740 1840	3+	N23 E08	3h 02m 1h 42m	. 18/2000-19/0400	Be, DLP, H, Kh, PS	во
19.	Sept. 21	1700	18	63	5.1	1330	1335	3	N10 W06	3h 30m	21/1200-21/2115	B, B2, CJM, HG, HNS,	PO 11 No 1/11
1 27.	Sept. 21	1,00		03	J•±	1330	1337	3	M10 W00	3 30	21/1200-21/211)	H, K, OH, PS, RL, S	BO, M, Mo, WH
20.	Sept. 26	2100		31	2.0	1907	1952	3	N22 E15	1 <sup>h</sup> 53 <sup>t.</sup>	26/2100-26/2315	B <sub>1</sub> , Be, Bo, HG, HNS, H, K, Kh, OH, PS, RL, S	во, и, wн
21.	Oct. 20	2100	22	64	7.8	1637	1642	3+	S26 W45	4h 23m	20/1700-21/1400	B <sub>1</sub> , Be, Bo, CJM, DLP, HG	BO, M, Mo, WH
						,		J			,	HNS, H, JP, K, Kh, OH, PS, RL, S	1
	1958											, ,	
22.	Feb. 10	0600	14	37	3.2	9/2108	2142	2+	S12 W14	8h 52m	10/0500-10/2400	B <sub>1</sub> , B <sub>2</sub> , Be, Bo, CJM,	BO, M, Mo, WH
												DLP, HG, HNS, K, Kh, OH, PS, RL, S	
23.	Mar. 23	1500	34	53	3.2	0947	1005	3+	<b>S14 E78</b>	5 <sup>h</sup> 13 <sup>m</sup>	23/1500-23/1830	<u>B1</u> , B2, B0	BO, M, Mo, WH
24.	Mar. 25	1530	13	122	10.0	0557		2	S15 E50	9 <sup>h</sup> 33 <sup>m</sup>	25/0100-25/1545	B <sub>1</sub> , B <sub>2</sub> , Be, CJM, DLP,	BO, M, Mo, WH
												HG, HNS, K, Kh, L, OH, PS, RL, S	
25.	June 06	0600			Weak	0436	0448	3(2)	N16 W78	1 <sup>h</sup> 12 <sup>m</sup>	04/2300-06/1345	Be, Bo, DLP, HG, HNS,	BO, M, WH
26.	T-1 07	23.26	•							h =		<u>KOh</u> , Off, PS, S	].
۵۰.	July 07	0130	22	96	23.7	0020	Cilo	3+	N25 W08	Oµ 50 <u>m</u>	07/0100-07/600	B <sub>1</sub> , B <sub>2</sub> , Be, Bo, CJM, DLP, HG, HNS, JP, K,	BO, M, Mo, WH
27.	July 29	0400		30	1.5	0350	ozoli		cale take	ı <sup>h</sup> oı <sup>m</sup>	So Johan So Jaran	Kh, L, OH, PS, RL, S	
-1.	anth sa	0400		30	1.5	0259	0304	3	S14 W44	1. 01	29/0400-29/0500	B <sub>2</sub> , <u>Be</u> , Bo, DLP, HG, HNS, JP, K, <u>Kh</u> ,OH, RL,S	BO, M, Mo, WH
28.	Aug. 16	0600	16 (2 đb/hr		<b>&gt;</b> 15.0	0433	0440	3+	S14 W50	$r_{\mu}$ $s_{8_{m}}$	29/0600-16/1200	B <sub>1</sub> , B <sub>2</sub> , Be, Bo, CJM,	BO, M, Mo, WH
			(2 40).11	. ,								DLP, RG, HNS, JP, K, Kh, L, OH, PS, RL, S, EIV	
29.	Aug. 21	1400	0.2	19	3.0	20/0042	0045	3(2+)	N16 E17	13 <sup>h</sup> 18 <sup>m</sup>	21/1400-21/1730	B <sub>1</sub> , <u>Be</u> , DLP, HNS, K, <u>Kh</u> ,	שט א טע
		2,00	db/hr	-/	500	20,5012	00+)	3(2.)	NIO EL:	13 10	21/1400-21/1/30	<u>L</u> , RL, S	50, F, Wh
30.	Aug. 22	1530	11	84	10.6	1417	1450	3	N18 W10	$1_{\mu} 53_{\mu}$	22/1500-22/1745	B <sub>1</sub> , B <sub>2</sub> , Bo, CJM, DLP, HG, HMS, K, Kh, L, OH, PS, RL, S	BO, M, Mo, WH
												PS, RL, S	
31.	Aug. 26	0100	1 db/hr	89	16.6	0005	0027	3	N20 W54	0 <sup>h</sup> 55 <sup>m</sup>	26/0100-26/0400	B <sub>1</sub> , B <sub>2</sub> , Bo, CJM, DLP, HG HNS, JP, <u>K</u> , <u>Kh</u> , <u>L</u> , OH,	BO, M, Mo, WH
												PS, <u>RL</u> , S	
32.	Sept. 22	1400	22	80	5.0	0738	0750	2 <b>+(</b> 2)	S19 W42	6 <sub>µ</sub> 55 <sub>w</sub>	22/0530-22/1730	B <sub>1</sub> , B <sub>2</sub> , Be, Bo, HG, HNS, JF, K, Kh, OH, PS, RL, S	BO, M, Mo, WH
	1959											,,,,,,	
33-	Feb. 13	0800	12	74	2.6	12/2301	2325	3+(3)	N13 E48	8h 59m	13/0800-13/1400	B <sub>1</sub> , Be, JC, JP	ВО
34.	May 10	2300	3 db/hr	170	22.0	2055	2140	3+	N19 E47		10/2300-11/0300	B1, B2, Be, CJM, DH, DLF	1
									•			EHO, JP, K, Kh, L, OH, R S, SL	
												_	

#### PCA'S WITH ASSOCIATED FLARES AND SOURCES (cont.)

No.	Date	PCA Beg.	Rise Time,hrs.	Duration hrs.	Abs.	Flar Onset	e Max	Imp.	Position	t	Range Start Times	Original Sources	Others
35.	July 10	0400	0.9 db/hr	360	20.0	0206	0230	3+	N20 E60	1h 54m	9/2000~10/1000	B <sub>1</sub> , B <sub>2</sub> , Be, CJM, DLP, <u>EMO</u> , K, <u>Kh</u> , L, OH, RL, S, SL	BO, M, Mo, WH
36.	July 14	0445	27	72	23.7	0325	0349 0527	3+	N17 E04	1 <sup>h</sup> 20 <sup>m</sup>	14/0445-14/0800	B <sub>1</sub> , B <sub>2</sub> , CJM, DLP, EHO, <u>K</u> , Kh, L, OH, RL, S,SL	BO, M, Mo, WH
37.	July 16	<b>&lt;225</b> 0	10	168	21.2	2114	2128	3+	N16 W30	1 <sup>h</sup> 36 <sup>m</sup>	16/2200-17/0600	B <sub>1</sub> , B <sub>2</sub> , CJM, <u>DH</u> , EHO, JP, <u>K</u> , Kh, <u>L</u> , OH, <u>RL</u> , S, <u>SL</u>	BO, Mo, WH
38.	Aug. 18	1100		60	3.0	1014	1030	3+(3)	N12 W33	oh 46™	18/1100-19/10xx	B <sub>1</sub> , <u>Be</u> , DLP, JC, JP, K, S	BO
İ	1960									h m			
39.	Mar. 29	0800	50	73	2.6	0640	0710	3(2+)	N13 E30	1 <sup>h</sup> 20 <sup>m</sup>	29/0800-29/1100	B <sub>1</sub> , B <sub>2</sub> , G	BO, Mo
40.	Mar. 30	2000		>36	5.0	1455	1540	3+(2)	N12 E13	5 <sup>h</sup> 05 <sup>m</sup>	30/1100-31/ <b>&lt;</b> 0700	B2, DH, EHO, GM, JP, K, I, S, Sat.	BO, M, Mo, WH
41.	Apr. 01	1000	6	73	3.6	0843	0859	3	NIS AII	1 <sup>h</sup> 17 <sup>m</sup>	01/0930-01/1000	B <sub>1</sub> , B <sub>2</sub> , <u>G</u> , GM, K, L, <u>S</u> , Sat	BO, M, Mo, WH
42.	Apr. 05	0400	>16	55	3.1	0215	0245	3(5)	N12 W62	1 <sup>h</sup> 45 <sup>m</sup>	05/0400-05/1400	B <sub>1</sub> , B <sub>2</sub> , EMC, <u>3</u> , DM, K, 1, <u>S</u> , Sat.	BO, M, Mo, WH
43.	Apr. 28	0230	12	30	3.5	0130	0137	3	S05 E34	1 <sup>h</sup> O <sup>m</sup>	28/0200-28/1000	B <sub>1</sub> , B <sub>2</sub> , EHO, G, GM, JF, K, L, S, Sat	BO, M, Mo, WH
44.	Apr. 29	0400	0.4 db/hr	114	14.0	0107	0210 0359 0554	3(2+)	N14 W21	2h 53 <sup>m</sup>	29/0200-29/700	B <sub>1</sub> , B <sub>2</sub> , <u>EHO</u> , G, GAY. JP, K, L, S, Sat	BO, M, Mo, WH
45.	May 04	1030	3.2 db/hr	8	3.4	1000	1016	3	N13 W90	oh 30 <sup>m</sup>	04/1030-04/1200	$\frac{B_1}{L}$ , $\frac{B_2}{S}$ , EHO, G, JP, K,	BO, M, Mo, WH
46.	May 06	1600	0.15-0. db/hr	33 103	16.0	1404	1448	3+	S09 E07	1 <sup>h</sup> 56 <sup>m</sup>	06/1400-06/1800	$B_1$ , $B_2$ , EHO, $\underline{G}$ , JP, K, L, S	BO, M, Mo, WH
47.	May 13	0620	0.7 db/hr	65	4.5	0519	0532	3+(3)	N29 W67	1h o1m	13/0620-13/0800	$B_1$ , $B_2$ , $EHO$ , $G$ , $JP$ , $\underline{K}$ , $\underline{L}$ , $\underline{S}$ , $Sat$	BO, M., Mo, WH
48.	Sept. 03	0500	31	89	2.7	<u>0037</u>	0108	3(5+)	N18 E88	7µ 53m	03/0500-03/2300	$\frac{B_1}{S}$ , $\frac{B_2}{S}$ , EHO, G, JC, K,	BO, M, No, WH
49.	Nov. 12	1400	16	73	21.2	1315	1330	3+	N27 WO4	on 45m	12/1400-12/1600	$\frac{B_1}{Sat}$ , $\frac{B_2}{S}$ , $\frac{G}{S}$ , $\frac{JP}{S}$ , $\frac{K}{S}$ ,	BO, M, Mo, WH
50.	Nov. 15	0430	15	79	20.0	0207	0551	3+(3)	n26 w35	Sp 53 <sub>w</sub>	15/043-15/1200	B <sub>1</sub> , B <sub>2</sub> , G, JP, K, S, Sat.	BO, M, Mo, WH
51.	Nev. 21	0200	15	51	3.0	20/1 <u>955</u> 20/ <u>2114</u>	2020 2135	3(1) 3(2)	N25 W90 N28 W90	6h 05 <sup>m</sup> 4h 46 <sup>m</sup>	21/0000-21/1300	<u>B</u> 1, B₂, G, JP, S	BO, M, Mo, WH
	1361							_		.h · F			
52.	July 11	2200	0.08 db/hr		1.3	1615	1659 1710	3	S07 E31	5h 45 <sup>m</sup>	11/2200-11/2400	<u>L</u> , M, Sat	BO, M, Mo
53.	July 12	1900	0.8 db/hr	72	17.0	0950	1025	3+(3)	S07 E22	9 <sup>h</sup> 10 <sup>m</sup>	12/1300-12/2115	B <sub>1</sub> , JP, L <sub>2</sub> , Sat	BO, M, Mo, WH
54.	July 18	1130	8	55	10.0	0920	1005	3+	S07 W59	$s_{\mu}$ 10 $_{\mu}$	18/1130-18/1200	B1, JP, L2, Sat	BO, M, Mo, WH
55.	July 20	2200			Weak	1553 1633 1828	1600 1653 1847	3+(3)	S05 W90	6h 07m 5h 27m 3h 32m		<u>le</u> , Sat	BO, M, Mo, WH
56.	Sept. 10	2000	18	79	2.9	1555	1610	1	N10 W90	4h 05m	10/2000-10/2300	B <sub>1</sub> , <u>BO</u> , Sat	M, Mo
57-	Sept. 28	2245	1.7 db/h pre-ssc 3.3 db/h post-ssc	1	3.3	2202	2223	3	N13 E29		10/2245-10/2335	B <sub>1</sub> , <u>L</u> , L <sub>2</sub> , <u>Sat</u> , <u>Bal</u>	BO, M, Mo, WH
	<u>1963</u>	al								<b>L</b>			
58.	Sept. 20	2400	15	54	3.1	2314	2403	2	N10 W09	oh 46 <sup>m</sup>		<u>B1</u> , BO,	
59•	Sept. 26	0745	8	89	4.6	0638	0717	3	N13 W78	1" 07"	26/0730-26/0745	B <sub>1</sub> , <u>BO</u>	L

TABLE A-2
SOURCES USED FOR POLAR CAP ABSORPTION DATA

A	Anderson	J. Geophys. Res.	<u>69</u>	1964	1743-1753
в	Railey	Planet. Space Sci.	12	1964	495-541
B <sub>2</sub>	Bailey	J. Phys. Soc. Japan	<u>17</u> A-1	1962	106-112
Be	Besprozvannaya	J. Phys. Soc. Japan	17 A-1	1962	146-150
Во	Bookin	J. Phys. Soc. Japan	<u>17</u> A-1	1962	150-151
во	Basler & Owren	U. Alaska Geophys. Inst.	R-152	1962	189 pp.
СЛМ	Collins, Jelly & Matthews	Can. J. Phys.	<u>39</u>	1961	35 <b>-52</b>
DH	Dodson & Hedeman	Ark. Geofysik	<u>3</u>	1962	469-470
DLP	Dvoryashin, Levitskii & Pankratov	Soviet Astron. A.J.	<u>5</u>	1961	311-325
EHO	Egeland, Hultqvist & Ortner	Ark. Geofysik	3	1962	481-488
FW	Freier & Webber	J. Geophys. Res.	<u>68</u>	1963	1605 <b>-16</b> 29
Go	Gosling	J. Geophys. Res.	<u>69</u>	1964	1233-1238
G	Gregory	J. Geophys. Res.	<u>68</u>	1963	3097-3107
GM.	Greenstadt & Moreton	J. Geophys. Res.	<u>67</u>	1962	3299-3316
Н	H511	J. Phys. Sec. Japan	<u>17</u> A-1	1962	97-102
HS	Hakura & Goh	J. Radio Res. Lab. Japan	<u> 5</u>	1959	635-650
HNS	Hakura, Nagai & Sano	Rep. Ionosph. Space	<u>15</u>	1961	14-30
JC	Jelly & Collins	Res. Japan Can. J. Phys.	<u>40</u>	1962	706-718
JP	Jenkins & Paghis	Can. J. Phys.	<u>41</u>	1963	1056-1075
ĸ	Kahle	U. Alaska Geophys. Inst.	R-129	1962	76 pp.
Kh	Khocholava	Geomag. Aeronomy.	<u>2</u>	1962	90-96, 907+913
			3	1963	735-740
Ll	Leinbach	U. Alaska Geophys. Inst.	R-127	1962	230 pp.
$\mathbf{L}_2$	Leinbach	U. Alaska Geophys. Inst.	R-126	1962	16 pp.
H	Malitson	NASA TR	R-169	1963	109-117
NW.	Malitson & Webber	NASA TR	R-169	1963	1-17
Мо	Modisette	Manned Spacecraft Center Eng. Des. & Oper, Ed.Purs et.al.	er,	1964	97-104
ОН	Obayashi & Hakura	J. Geophys. Res.	<u>65</u>	1960	3143-3148
PS	Piggott & Shapley	Antarctica Res. Geophys. Mon.	<u>7</u>	1962	111-120
RL	Reid & Leinbach	J. Geophys. Res.	<u>64</u>	1959	1801-1805
S	Sinno	J. Geomag. Geoelect.	<u>13</u>	1961	1-10
SL	Snapley & Lincoln	Ann. IGY	<u>16</u>	1962	289 pp.
WH	Warwick & Haurwitz	J. Geophys. Res.	<u>67</u>	1962	1317-1332

						SPALL	PCA'S	REPORT	ED BY TWO	OR MORE INDEPENDENT	OBSERVERS				
		PCA	Rise	Duration		Flar Onset			Transition	PCA Range Start Times	Original Sources	T	Sunspot		•
No.	Date 1955	Be⁻.	Time, hrs.	hrs.	db.	onsec	Max.	imp.	F081:10H	Range Start Times	30th Ces	Flage	Durispos		
Sì	1-16	1600		48	2					16/1600-16/2230	B <sub>2</sub> , Be, JC	3065	11218	61	wa6
31	1256	1000		40	٤	•,				10/1000-10/2230	52, 56, <u>50</u>	3007	11210	01	<b>N3</b> 6
S2 .	4-27	2000		48		1546		1+	S14 E14	27/2000-27/2200	Be, JC	3477	11596	324	<b>S1</b> 5
-	1957	2000		45		-/**		-		-1,2500 2.,7230	, <u></u>			٠.٠	<b>51</b> ,
<b>S</b> 3	2-21	1800		72-96		1605	1930	3+	N20 W30	21/1800-22/1600	Be, DLP, JC	3856	12140	10	<b>N</b> 25
-5		2000					- /50			1,	л <del>р</del> , s	30,72			
54	4-06	0800	12	66	3.2	05/1433		1	S15 W90	and the second	B <sub>1</sub>	3907	12235	513	S15
S:	4-12	1700				11/1722	1738	3(2+)	\$23 E04	11/1300-12/1700	Be, <u>JC</u> , <u>JP</u>	3923	12254 12258	43	\$22
s:	7-01	1200				r				01/ 000-01/1200	<u>сум</u> , н				
s-	7-28	1500		24	Weak	1340	1402	2	524 WB3	28/1500-28/2100	н, <u>JC</u> , <u>PS</u>	4070	12496	13%	S23
<b>3</b> 3	2-22	1000				0643		1+	N24 W32	22/1000-22/1200	н, рѕ	4151	12622	85	N23
						0732	0750	2	N23 W38		_				~
sn.	11-05	0030	10	46	2.6	04/0058 1058 1732	0102 1740	1	520 W38 524 W39 525 W45	04/2300-05/0300	B <sub>1</sub> , Be, DLP, H JC, <u>PS</u>	4207	12732	240	524
sin	12-17	1300			Weak	0734	0737	2	N20 E41	17/0300-17/1600	н, <u>јр</u> , РЅ	4314	12855	313	N18
311	12-29	2300		30	Weak	2229	2230	5	N25 ¥50	28/2300-28/2400	н, <u>JC</u> , <u>PS</u>	4321	12874	263	N22
	1258														
512	3-11	2400			Weak	0030	0042	1	N11 E02	11/0300-11/1000	Be, <u>JC</u> , PS, S	4449	13076	307	N11
S13	3-14	2200				1454	1507	2	S21 W84	14/1500-14/2200	Be, CJM, JP_PS,S	կկկե	13063	15	S18
S14	3-31	1600				31/2005	0014	2	S17 W22	31/NTG -31/160C	JC, JP		(13103)	32	\$13
	-				:	30/2345 31/0038 31/0025	2347 0052 0031	1 2 2	S10 W31 S08 W23 W37 E59			4476 4484	13110 13110 13118		
S15	4-10	1990	6	68	4.4	f					B <sub>1</sub> , B <sub>2</sub> , Be, CJM JP, K, L <sub>1</sub> , PS, R	 L,s 			
	1259														
516	1-26	1500				0842	0900		N16 W61		JC, JP	4069	13878	106	N17
817 318	6-13	1330			1.5	0357	1358		N17 E58	13/0900-13/1330	Be, JC, K	5204	14211	330	N17
JI:	3+05	0490		48	small	1923 1648 1947 (92/9310	1738 1704 1953 0434	2+ 2+ 1+ 1+	N12 E60 S12 W52 N09 W15 N17 E21		<u>B</u> 2, €HO, C	5354 5340 5344 5340	14414	355	N12
	<u>1960</u>														
31.	1-11	2200		76	A2	2040	2124	3	N22 E03	11/2200+12/0700 13/1600-13/2000	$B_1, G \\ JC, \overline{S}$	5527	14660	101	N1
520	9-26	0300		120	S qp	0525	0537	1+	S22 W64	25/2100-26/2300	G, JC, <u>JP</u> , K	5858	15043	353	<b>S1</b> 9
_	1961														
521	11-10	1500			2 <b>d</b> b	1434	1444	1+	N19 W90	10/1500-10/1600	B <sub>1</sub> , Ba, BO	6264	15461	5	NO
	1962				-										
:55	2-01	20 <b>3</b> 0			1-2db	0901	0907	2	N10 M36		BO, Sat	6326	15507	298	N1C
	1263					,									
23	4-15	1200				1034	1125	2	S11 W06		BO, Ma	6766	15714	246	S12

TABLE A-4
PCA'S REPORTED BY GREGORY OWLY DURING 1960

		PCA	W	<b></b>	Flare			T 8-3-	L		T.,				
No.	Date	UT Hrs	Duration Days	Beginning	Max	Importance	Position	Delay Hours	Source Station	Plage	Region	Mo.	Ł		Plage Age
### G2	1-15	03	1	1334		2	s20 W68	14	кіля	5525	1422-45	14657	121	\$17	New
G3	2-7	07	4	#06/1340 06/1426	1344	1 2	S13 W81 S15 W03	17.5	Scott B.	5551	1423-3	14701	196	S13	4
G4	2-15	10	4					1							1
G5	2-29	16	8	*1522 0140	1546	1 2	N22 E04 832 W56	<1	Scott B.	5586	1424-21	14738	172	N24	2, 5
Gó	3-10	18	2.5	1716	1720	1	<b>M25 E08</b>	1	HLS	5592	1424-39	14751	41	<b>N</b> 25	3
G7	3-17	18	3	1616	1620	1	#05 W32	1	HLS	5597	1425-1		349	N06	5
***	1														
G11	4-15	10	4	#0717 0950	0719 0957	1	N12 W17 N23 B68	2.5	HLS	5627	1426-11	14796	313	N11	New
G16	5-09	08	> 3	0704	0734	3	S16 E55	1	HLS	5657	1427-11	14831	279	S09	3
G18	5-17	15	2	*0418	0425	1+	SO9 E33	11	HILS	5663	1427-28	14840	197	S12	3
G19	5-26	10	3	0818	1418 0928	1- 2+	N18 W08 N16 W15	2	HLS	r660	1427-40	14849	206		1
620	6-01	14	6	0824	0900	3+	M28 E39	6	HLS	5669	1428-02	14867	126 343	N13	2
G21	6-15	10	2	0635	0653	2	S09 E08	1	HLS	5695	1428-22	14888	199	N30 S12	1
G22	6-25	17	> 2	1131	1215	3	N19 EO4	6	HILS)	""	140-00	14000	-77	<b>511</b>	
G23	6-27	23	>1	2140	2156	3	W17 W23	1,5	HIS	5713	1428-39	14908	69	N20	Nev
G24	6-28	19	1.5	#1214 1815	1217	1	#21 W37 #08 £68	7	HLE	,	2.22.37	21,02	٠,		
G25	8-11	24	5	1916	1929	2+	N22 E27	5	HLS	5794	1430-13	14981	143	<b>N</b> 20	2
G26	8-26	10	5	0847	0852	1	N17 W90	1	HLS	5802	1430-23	14989	68	N17	3
****											- 5- 5-	2.,,0,	•		]
***	İ							1							
<b>G2</b> 9	10-3	16	10	**				1	HLS						
G30	10-29	12	8	1026	1030	3	N14 E25	1.3	HLS	5909	1433-19	15099	185	N13	Nev
G31	11-10	18	>1	1009	1023	3	N28 E28	8	,			•	-		
G32	11-11	04	>1	0305	0340	2	N28 E12	1	HLS						
	11-14	22	>1	2114	2120	1+	N35 W27	1	A11	5925	1433-39	15114	28	N27	3
****								·	\						
G36	11 <b>-</b> 19	12	>2	1001	1059	1	N25 W90	2	Southern						
G38	12-06	05	6	5/1825	1839	3+	N26 E68	11	HLS	5959	1434-38	15151	8	N25	4
***	H No PCA H See H For	's Report For: Table 7:	ed by Gree G1, 28 S19, 20 G÷8, 9	gory and Othe	, 14, 15,	17, 27, 33, 47, 48, 49,							-	,	